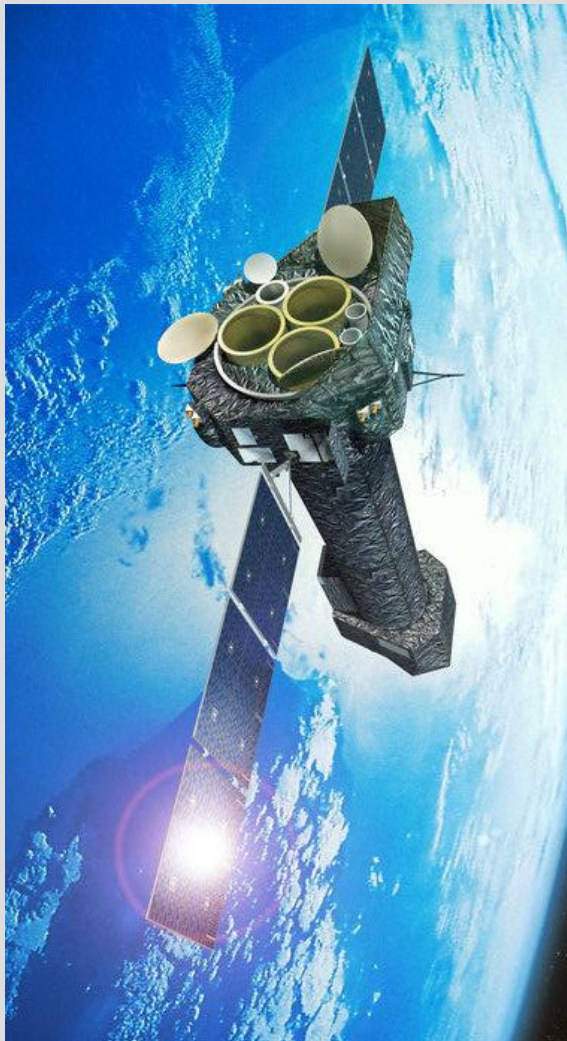


The Complete XMM-SERVS Survey: A Sensitive X-ray Survey of the LSST Deep-Drilling Fields

W.N. Brandt (Penn State), Q. Ni (Penn State), for the XMM-SERVS Team



Qingling Ni

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Large Effective Area ($\sim 2000 \text{ cm}^2$)

Good Field of View ($\sim 800 \text{ arcmin}^2$)

Good Angular Resolution

Broad Bandpass (0.3-10 keV)

Healthy – hopefully, 10+ more years

Rubin Observatory

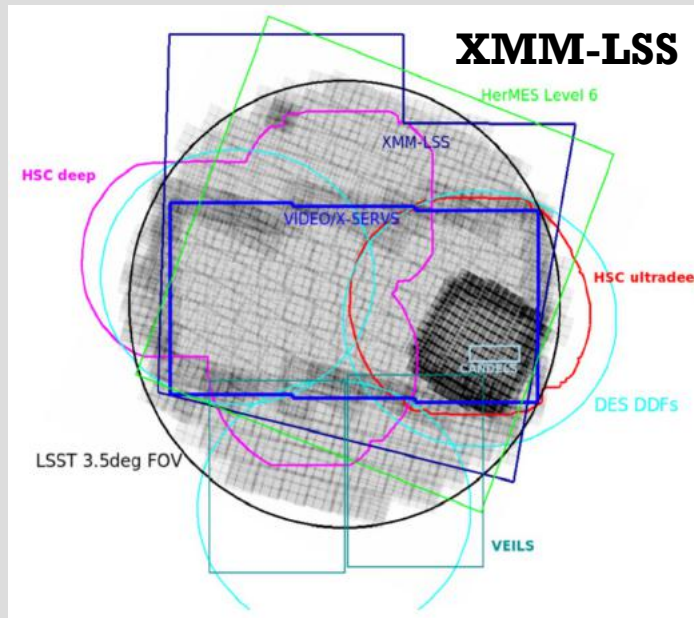
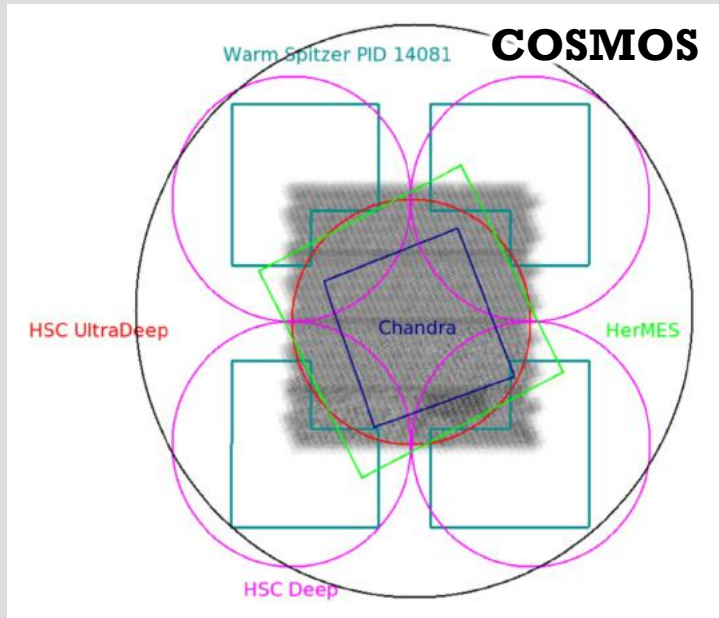
Will conduct Legacy Survey of Space and Time (LSST)

Main survey over $\sim 18000 \text{ deg}^2$

Also “mini-surveys” for $\sim 10\%$ of time

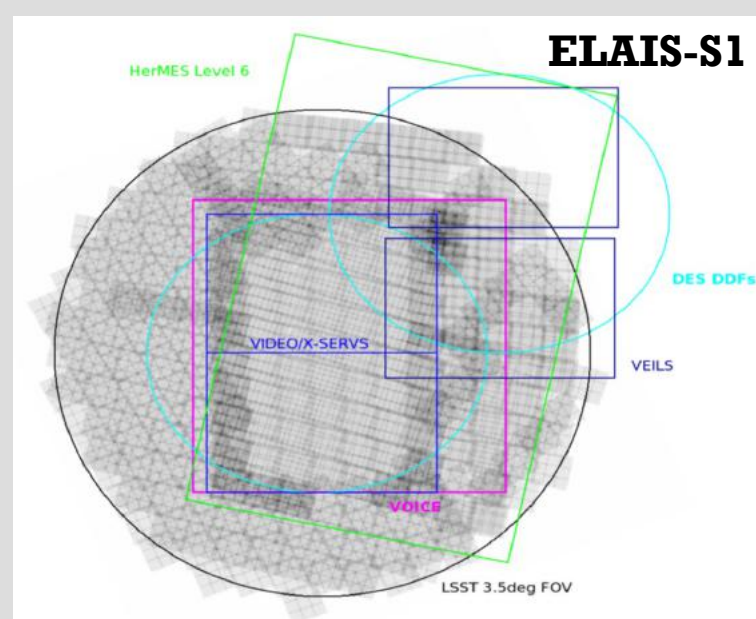
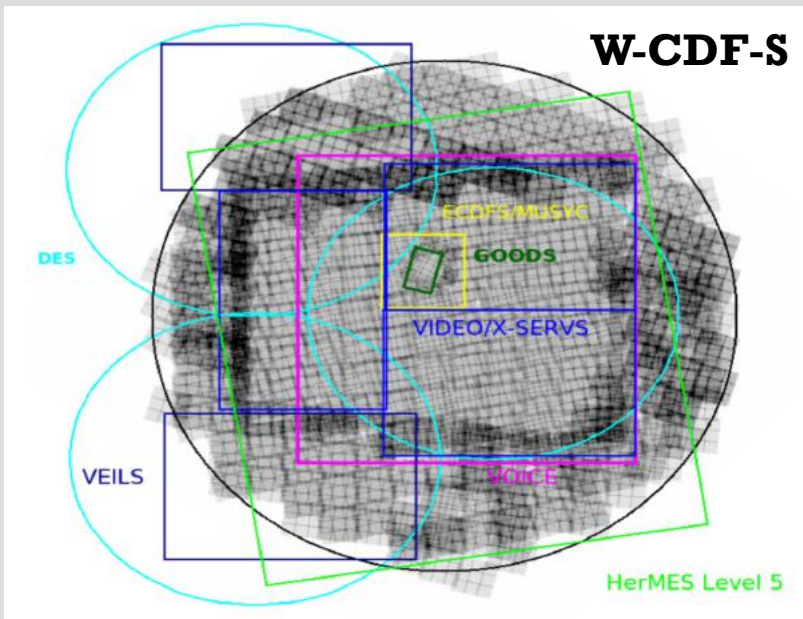


Rubin's LSST Deep-Drilling Fields



Each LSST field is 10 deg² (large black circles).

Likely will observe every 2-3 nights in *grizy* when visible for 10 yr (also *u*).

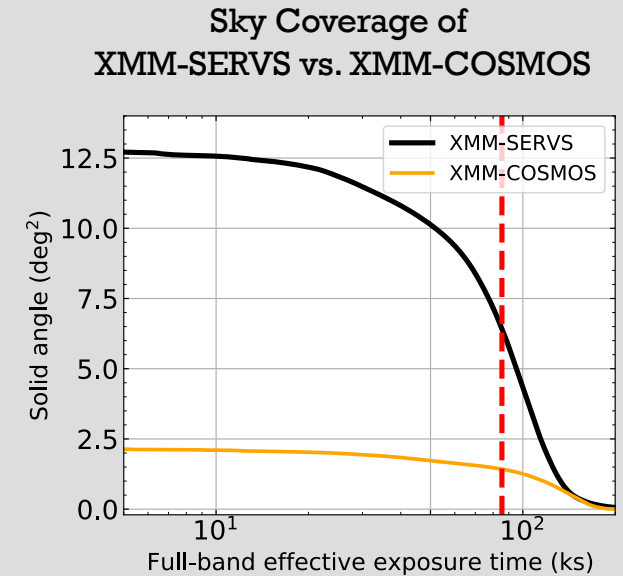
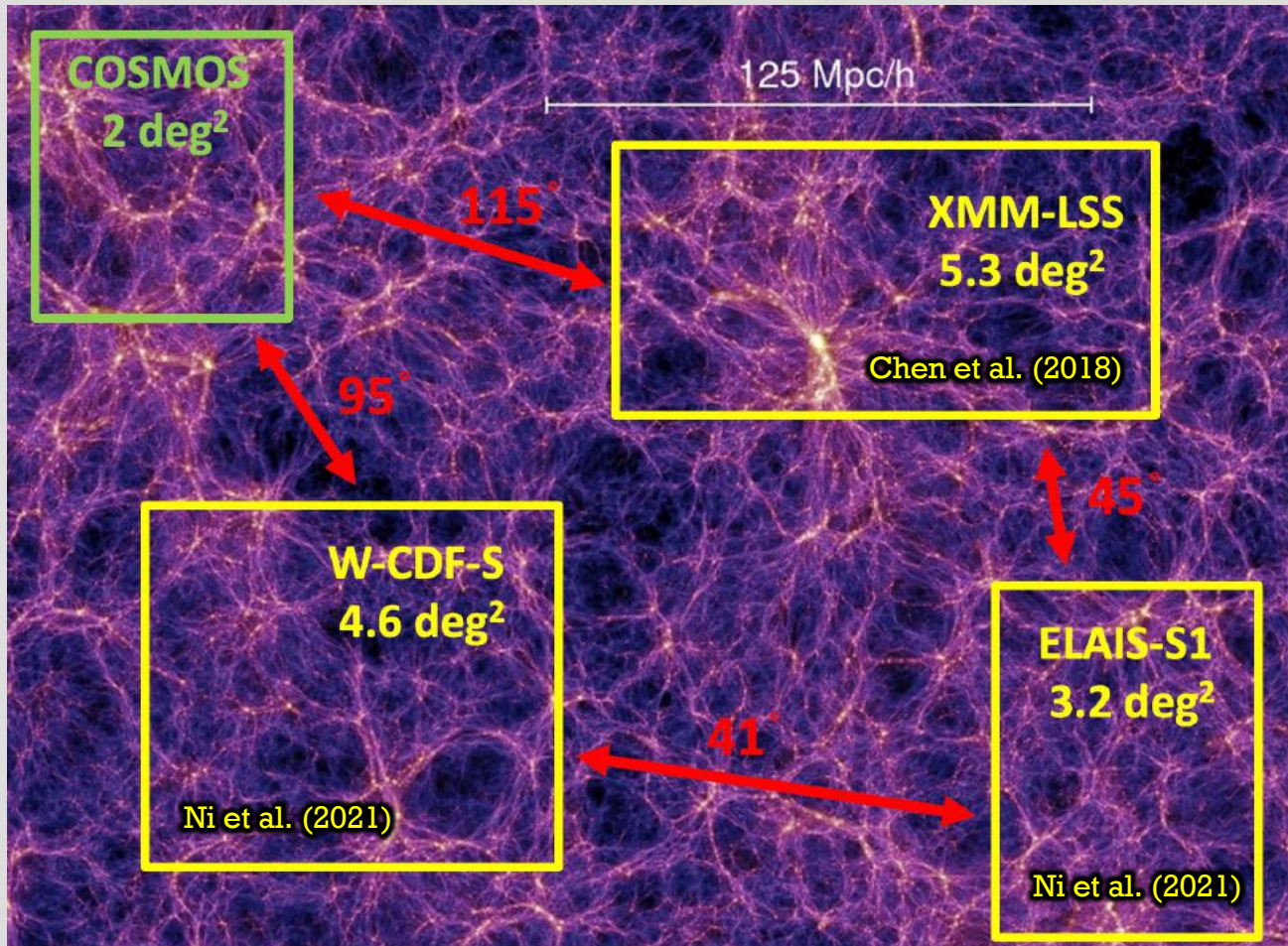


The Spitzer Extragalactic Representative Volume Survey (SERVS) and the DeepDrill survey



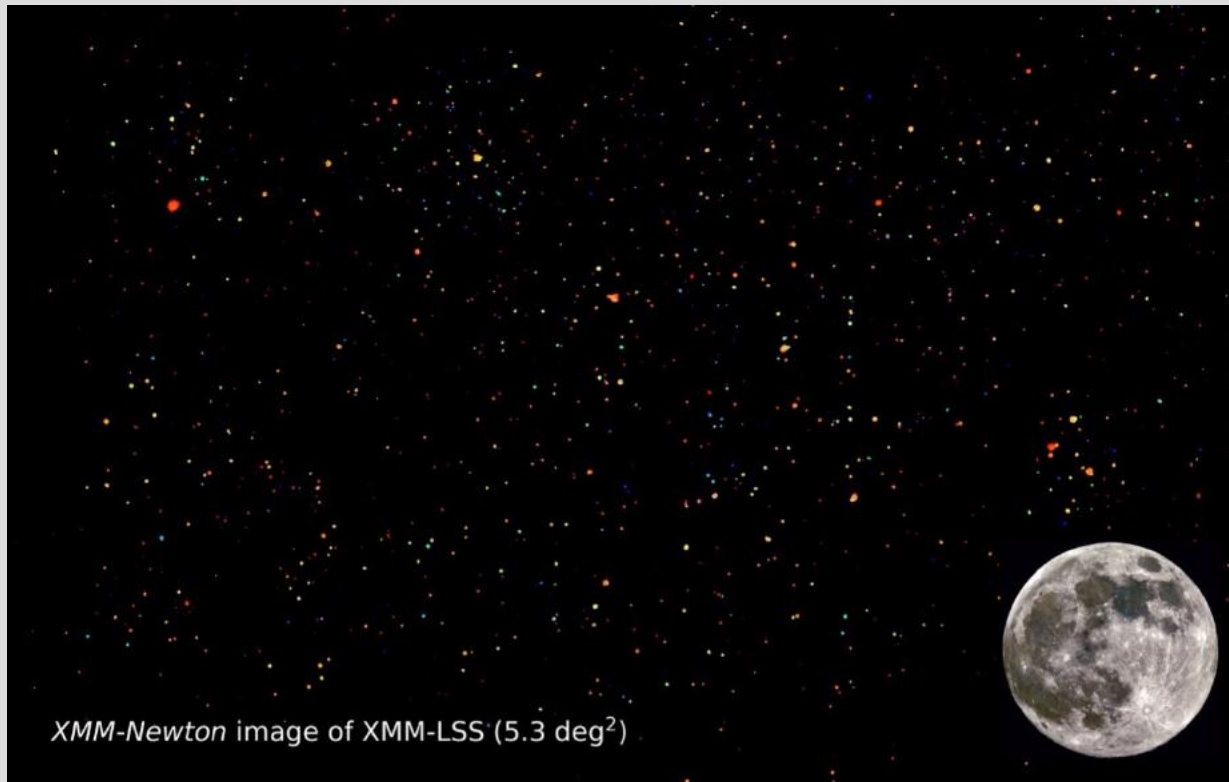
Brandt et al. (2018)
arXiv:1811.06542

The 5 Ms XMM-SERVS Heritage Program

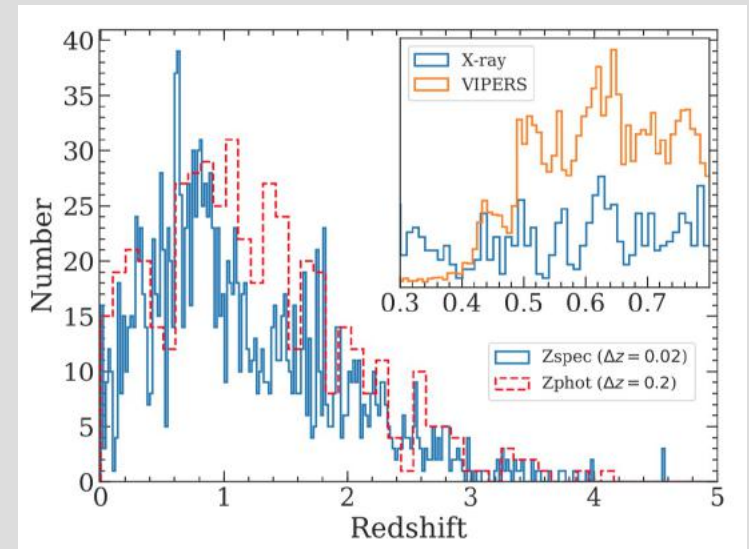


At uniform 50 ks XMM-Newton depth, detect ~ 10,200 AGNs and many X-ray groups/clusters.
SMBH growth across the full range of cosmic environments and SMBH/galaxy connections.
Great legacy value as LSST/DES DDFs and MOONS/PFS, ToITeC/ALMA, & MOS-RM fields.
Ground-truth AGN sample for calibrating LSST AGN selection in DDFs and main survey.

Results from the XMM-LSS Field



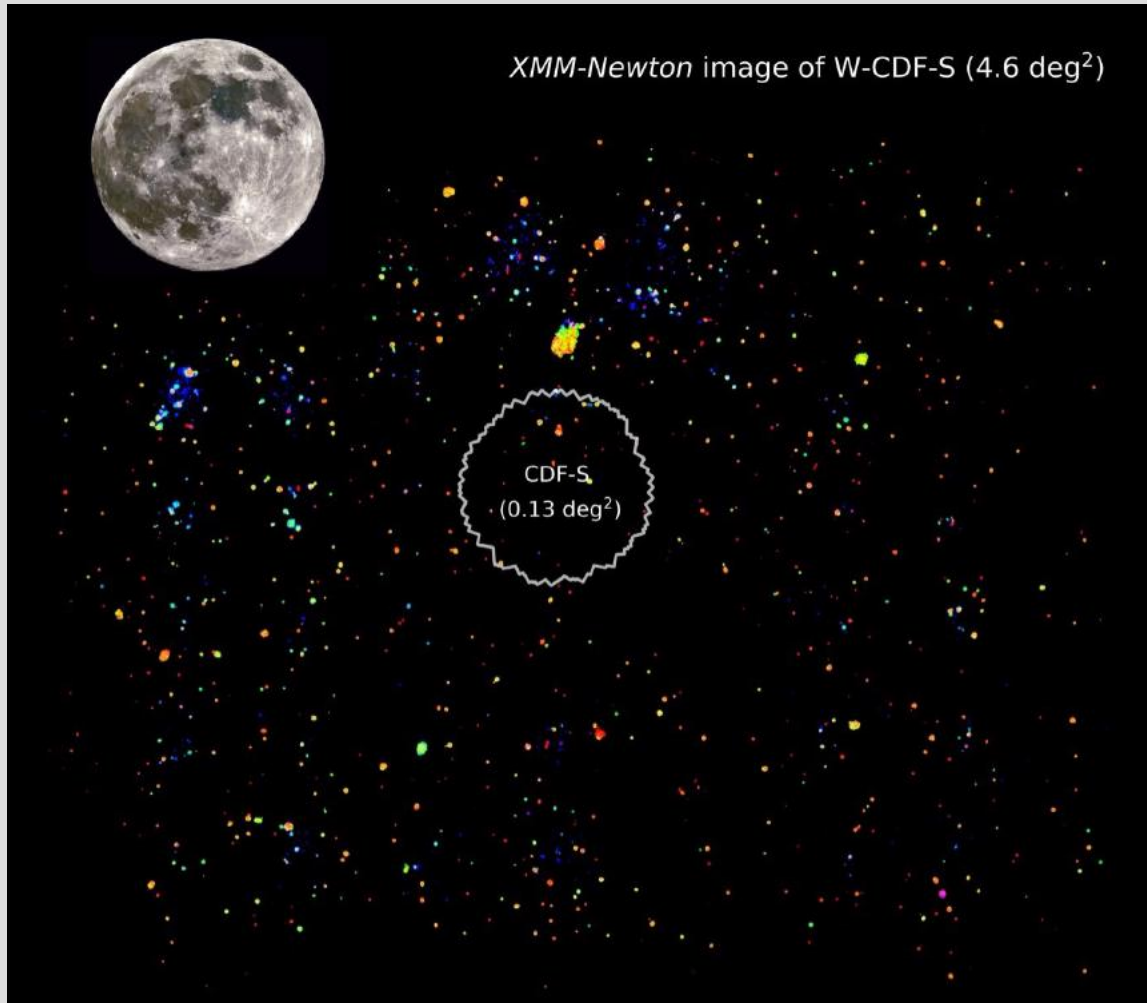
Chen et al. (2018)



5242 total X-ray sources - 90% good identification rate (assessed with Chandra)

In our prime 4.5 deg², more than 70% of sources have spectroscopic or high-quality photometric redshifts.

W-CDF-S Now Submitted!



Ni et al., submitted

80 observations done with total good exposure of 1.8 Ms, covering 4.6 deg^2 .

Background flaring fine overall; used re-observations to fill “holes” for exposure uniformity.

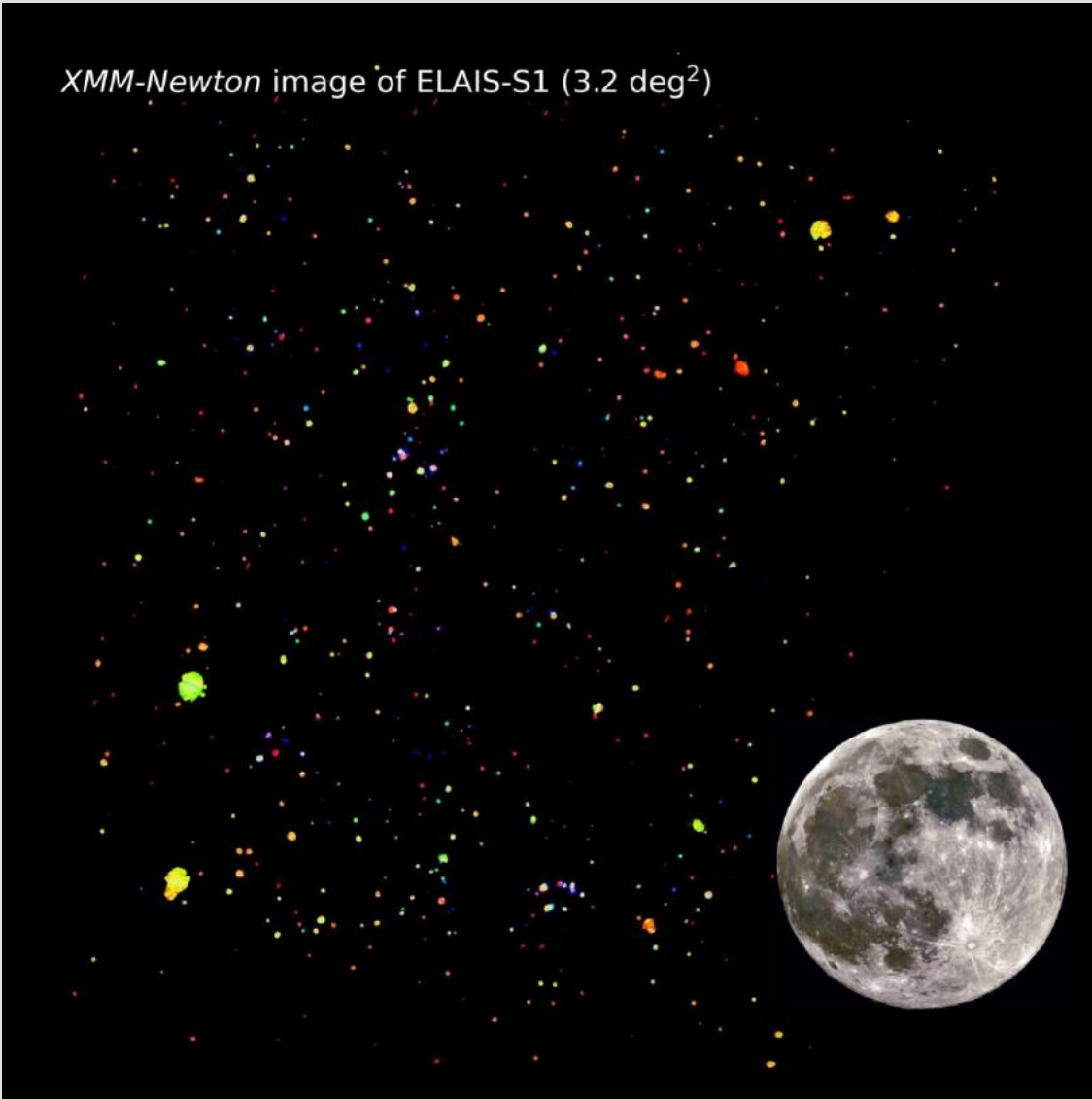
4053 X-ray sources detected, and almost all are new!

89% have reliable multiwavelength counterparts in Spitzer/VIDEO/HSC.

68% have spec-z or high-quality photo-z (83% in prime multi- λ area).

ELAIS-S1 Now Submitted!

XMM-Newton image of ELAIS-S1 (3.2 deg²)



31 observations done with total good exposure of 0.9 Ms, covering 3.2 deg².

Modest background flaring.

2630 X-ray sources detected.

87% have reliable multiwavelength counterparts in Spitzer/VIDEO/DES.

65% have spec-z or high-quality photo-z (85% in prime multi- λ area).

Source Numbers Summary

For all three XMM-SERVS fields together, detect

11,900 X-ray point sources

10,200 AGNs (86%)

About 2800 Type 1 AGNs (27%)

About 7400 non-Type 1 AGNs (73%)











About 1250 galaxies and 450 stars

XMM-SERVS point-source sky density is $\sim 910 \text{ deg}^{-2}$
compared to $\sim 170 \text{ deg}^{-2}$ typically for eROSITA
(and XMM-Newton positions considerably better).

1.6 Million Photometric Redshifts

OPEN ACCESS

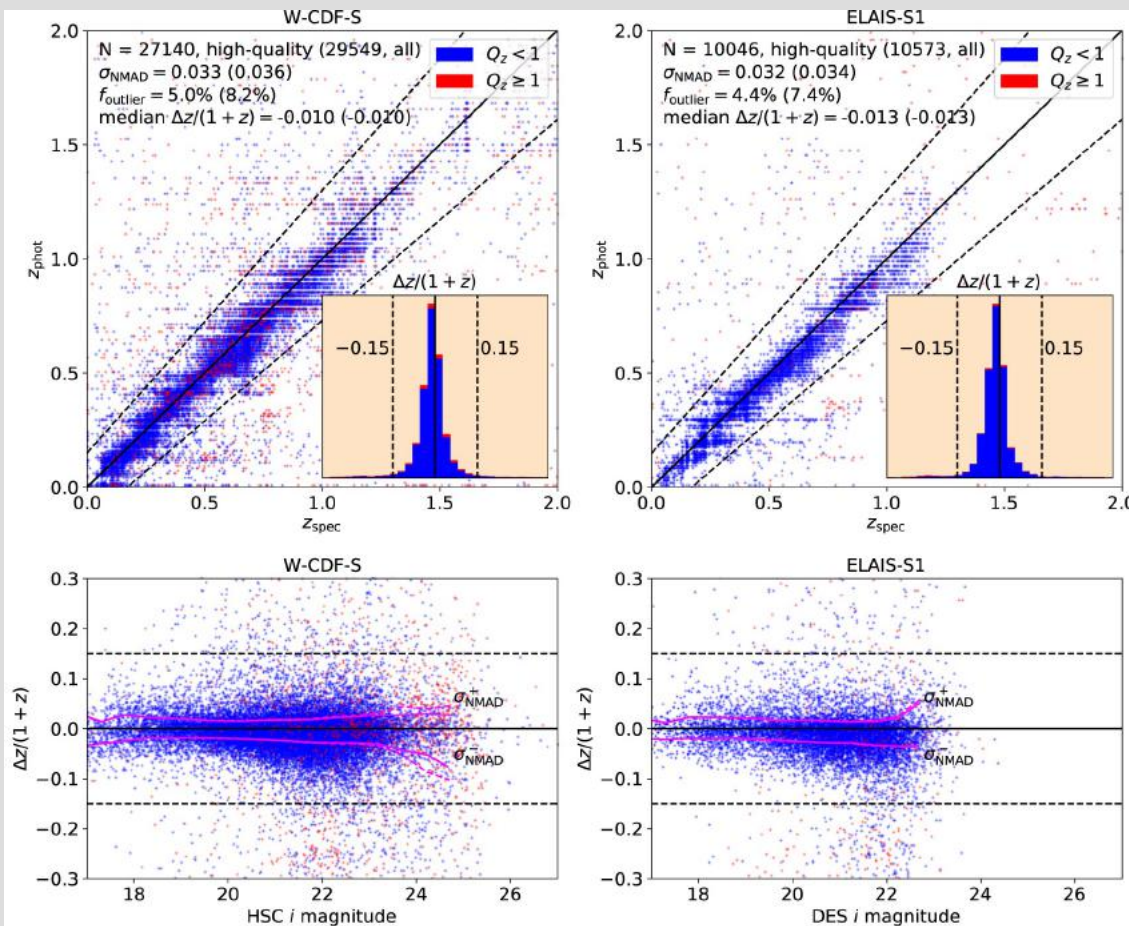
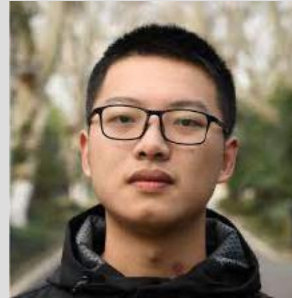
Photometric Redshifts in the W-CDF-S and ELAIS-S1 Fields Based on Forced Photometry from 0.36 to 4.5 Microns

Fan Zou^{1,2} , Guang Yang^{3,4} , W. N. Brandt^{1,2,5} , Qingling Ni^{1,2} , Franz E. Bauer^{6,7,8} , Giovanni Covone^{9,10,11} , Mark Lacy¹² , Nicola R. Napolitano⁹ , Kristina Nyland¹³ , Maurizio Paolillo^{9,10,11}  + [Show full author list](#)

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[Research Notes of the AAS, Volume 5, Number 3](#)

Citation Fan Zou et al 2021 Res. Notes AAS 5 56

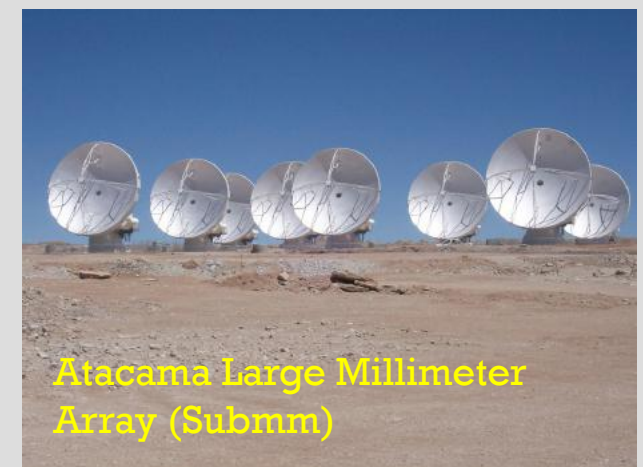
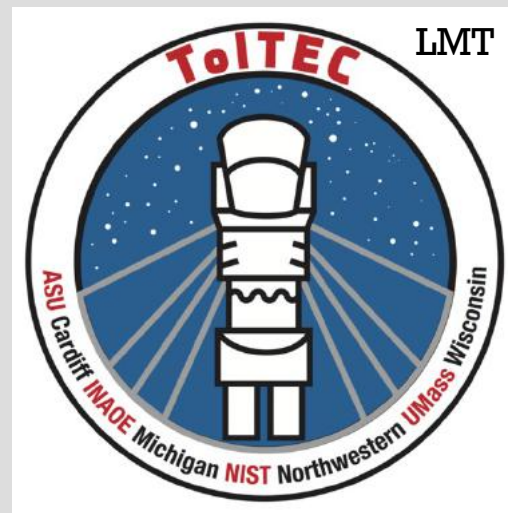
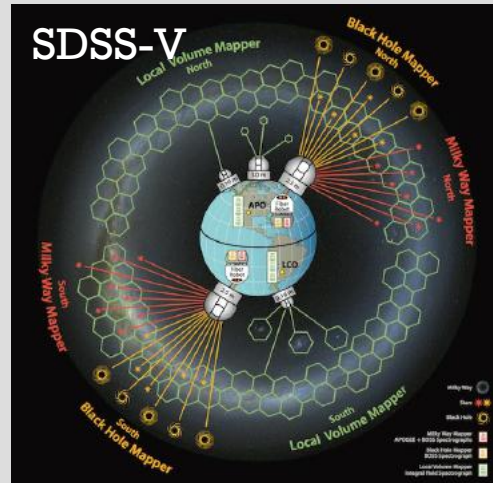


Also have delivered:

HSC imaging catalog for W-CDF-S - Ni et al. (2019)

Multi-band forced photometry catalogs for ELAIS-S1 and W-CDF-S – Zou et al. (2021) and Nyland et al. (in prep)

Complementary Multiwavelength Data Flooding In for XMM-SERVS!



The End

Public data release for W-CDF-S and ELAIS-S1 soon!

<http://personal.psu.edu/wnb3/xmmserve/xmmserve.html>

W-CDF-S XMM-Newton
false-color image

